Low- and mid-latitude ionosphere as observed by the Swarp by the Swarm constellation

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The Swarm constellation is the European Space Agency's mission for the Farth's state of the European Space Agency's of the European Space Agency Space Agenc mission for the Earth's magnetic field, which consists of the three satellites equipped identically. The constellation was launched in November 2013 into a polar (inclination angle~87.30) circular orbit. During the first two months of the commissioning phase all the three satellites were at the commissioning phase all the three satellites were at the same altitude (-500 km) in a pearls-on-a-string configuration. Afterwards, the satellites conducted orbit maneuver to be placed into separate orbits. Since April 1014 the two lower satellites (Swarm-Alpha and Swarm-Charlie) have been flying side-by-side with a Swarm-Charlie) have been flying side-by-side with a congitude separation of about 1.50. The upper satellite, warm-Braye shall be supported by the bigher warm-Braye shall be separated by the bigher warm-Braye shall be supported by the bigher w warm-Bravo orbiting on a higher altitude (-50 km higher than that of the lower pair) has a slightly different orbit clipation, and the clipation and the cli clination angle (~87.75o), by which the local time of the bital plane is separated gradually from those of the wer pair. The constellation frequently encounters plasma insity irregularities in the low- and mid-latitude topside topsphere. In this presentation we investigate selected ents in the low- and mid-latitude topside selected in the low- and mid-latitude topside to ents in the low-/mid-latitude ionosphere (plasma density pletion or enhancement), as observed by the Swarm nstellation during its first-year mission.

## 20 [111-2-4]

ospheric space weather monitoring by RMOSAT-3/COSMIC and following RMOSAT-7/COSMIC-2

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ne past decade, the paucity of ionospheric tions has made it almost impossible to reconstruct ee-dimensional structures of global ionospheric density. The FORMOSAT-3/COSMIC (F3/C) consists of six microsatellites, which has proven powerful tool in probing global high-resolution ric vertical electron density profiles by radio techniques. By accumulating ons, the three-dimensional structures of global ric electron density can possible be retrieved.

In the success of F3/C mission for atmospheric ospheric researches, the National Space ion in Taiwan has proposed a follow-on mission ormosa Satellite-7/Constellation Observing System eteorology, lonosphere and Climate-2 AT-7/COSMIC-2, hereafter shortened to F7/C2),

which deploys an operational constellation system which satellites to receive US GPS, Russian GLONAS twelve European Galileo Stated for deployment stands and European observations. Will further provide most operation of the F3/C occultation of th and European standard occultation observations. Stated to construct the standard occultation observations will further provide more standard occultation occultation sound to times the number of the F3/C occultation sound four times the number of and space weather monitor weather to recasting and a major advance in incompanion. tour times the number of and space weather monitoring for the future mission weather monitoring for the future mission weather monitoring for the future mission will be paper. presented in this paper.

09:40 [HI-2-5] Evaluation of IRI-2012 by comparison with Evaluation of and incoherent scatter radar JASON-1 TEC and incoherent scatter radar JASON-1 during the 2008-2009 solar observations during the 2008-2009 solar minimum period

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minimum solar 2008-2009 The 2008-2009 solar extended and therefore unprecedentedly deep and extended and therefore unprecedentedly deep for the data-driven more unprecedentedly dependent of the data-driven model is should be difficult to expect for the data-driven model is should be difficult to expect for the data-driven model is should be difficult to the International Heleichlock the ionospheric parameters during well reproduce the ionospheric parameters during the well reproduce the ionospheric parameters during the period. In this study, we compare the IRI-2012 period. TEC data from JASON-1 satellite and electron global TEC data from JASON-1 satellite and electron profiles observed from Millstone Hill Incompared to the ionospheric parameters during the period of the ionospheric parameters during the period of the ionospheric parameters during the ionosphe global TEC data from SASSIVE Satellite and electron density profiles observed from Millstone Hill Incoheren Scatter Radar (ISR) and EISCAT at middle and his latitudes, respectively. For the global mean ionosphere latitudes, respectively. daily mean TECs calculated from the model and data are daily mean TECs of other during the two-year period and compared with each other during the two-year period and it is found that IRI underestimates TEC by about 344 it is found that IRI underestimates TECs are produced. Global TEC maps for two TECs are produced in the magnetic latitude and magnetic local time coordinate which shows the overall underestimation of IRI except to the region around the equatorial anomaly only during the daytime, regardless of season. In the daytime equators region, IRI slightly overestimates TEC and this aspect region, IHI slightly overtextivity was also reported in the previous study (e.g. see Jee et al., 2005). In order to further investigate the differences at middle and high latitudes, which are responsible for the overal underestimation in the global mean ionosphere, compared the electron density profiles between IRI and ISRs at Millstone Hill and EISCAT for the middle and high latitudes. During the daytime, however, it is not ven clear that IRI underestimates TEC. In particular, the density profiles show large discrepancy in the topsor ionosphere, which requires further investigation since the ISR data may not be reliable in the topside ionosphere for low solar activity.

제3발표장 (육백)

■ III-3 달과학 | 좌장: 최영준 (천문연)

08:30 [III-3-1]

Lunar dust environment using LDEX data onboard LADEE

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